

[Lecture 4]

Use

- ⊙ Problem set 2, due Monday 2/25/02  
 Problems 33, 35, 37, 38, 39 From Sakurai chapter 5

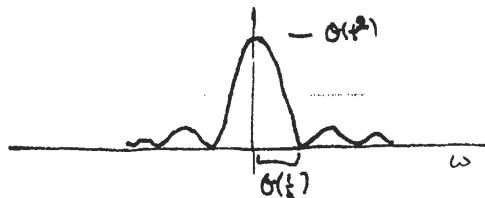
Last time:

1st order TDPT for harmonic (constant  $V(t)$ ).

For constant  $V(t) = \hat{V}$ ,  $C_n^{(1)} = \frac{\hat{V}_{ni}}{i\hbar\omega_{ni}} [1 - e^{i\omega_{ni}t}]$

$$P^{(1)}(i \rightarrow n) = \frac{4|\hat{V}_{ni}|^2}{(E_n - E_i)^2} \text{Si}^2 \left[ \frac{(E_n - E_i)t}{2\hbar} \right]$$

$$= |C_n^{(1)}(t)|^2$$



$\lim_{t \rightarrow \infty} \sim t\pi \delta(\omega)$

~~$\lim_{x \rightarrow \infty} \frac{\text{Si}^2(x)}{x^2} = \pi \delta(x)$~~

Transition rate  $W_{i \rightarrow n} = \frac{d}{dt} |C_n^{(1)}|^2 = \frac{2\pi}{\hbar} |\hat{V}_{ni}|^2 \delta(E_n - E_i)$

(t big enough for EC, short enough, TDPT ok.)

Introduce density of states  $\rho(E)$

$W_{i \rightarrow n} = \frac{2\pi}{\hbar} |\hat{V}_{ni}|^2 \rho(E_n) \Big|_{E_n = E_i}$

↑  
final states w/  $E_n = E_i$

Fermi's Golden Rule

$\left[ |\hat{V}|^2 = \lim_{\Delta E \rightarrow 0} \frac{1}{\Delta E} \int_{-0\epsilon/2}^{0\epsilon/2} |\hat{V}_{ni}|^2 dE_n \right]$  if  $\rho, \hat{V}$  smooth on families of final states

Back to harmonic perturbation

$$V(t) = V e^{i\omega t} + V^+ e^{-i\omega t}$$

$$C_n^{(1)} = \frac{1}{\hbar} \left[ \underbrace{\frac{1 - e^{i(\omega_{ni} + \omega)t}}{\omega_{ni} + \omega}}_{\text{peaked near } \omega = -\omega_{ni}} V_{ni} + \underbrace{\frac{1 - e^{i(\omega_{ni} - \omega)t}}{\omega_{ni} - \omega}}_{\text{peaked near } \omega = \omega_{ni}} V_{ni}^+ \right]$$

$\omega \equiv -\omega_{ni}$  : Stimulated emission

$$|C_n^{(1)}|^2 \approx \frac{4 |V_{ni}|^2}{\hbar^2 (\omega + \omega_{ni})^2} \sin^2 \left[ (\omega + \omega_{ni})^2 t / 2 \right]$$

Transition rate  $\rightarrow$  state w/ energy  $E_n$  at large  $t$

$$\omega_{i \rightarrow n} = \frac{2\pi}{\hbar} |V_{ni}|^2 \delta(E_n - E_i + \hbar\omega)$$

$$\omega_{i \rightarrow [n]} = \frac{2\pi}{\hbar} \overline{|V_{ni}|^2} \rho(E_n) \Big|_{E_n \approx E_i - \hbar\omega}$$

total emission rate 

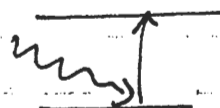
$\omega \equiv \omega_{ni}$  : absorption

$$|C_n^{(1)}|^2 \approx \frac{4 |V_{ni}^+|^2}{\hbar^2 (\omega - \omega_{ni})^2} \sin^2 \left[ (\omega - \omega_{ni})^2 t / 2 \right]$$

$$\omega_{i \rightarrow n} = \frac{2\pi}{\hbar} |V_{ni}^+|^2 \delta(E_n - E_i - \hbar\omega)$$

$$\omega_{i \rightarrow \text{cont}} = \frac{2\pi}{\hbar} |V_{ni}^+|^2 \rho(E_n) \Big|_{E_n = E_i + \hbar\omega}$$

total absorption rate



So - harmonic perturbation causes stimulated emission or absorption in units of  $\hbar\omega$ .

- Just what we expect if background made up of quanta w/  $E = \hbar\omega$ !

For transitions to occur & satisfy energy conservation, must have

(a) final states exist over continuous energy range, to match  $\Delta E = \hbar\omega$  for fixed perturbation frequency  $\omega$   
- or -

(b) Perturbation must cover sufficiently wide spectrum of  $\omega$  so that discrete transition with a fixed  $\Delta E = \hbar\omega$  is possible.

- Note that spectral lines are not really sharp, due to decay processes.

Note: For two discrete states,  $\omega_{i \rightarrow n}^{(abs)} = \omega_{n \rightarrow i}^{(em)}$  in semiclassical calc.  
since  $|V_{ni}|^2 = |V_{ni}^+|^2$

$\Rightarrow$  Detailed balance

[Really, only true @  $T = \infty$  when rad. field quantized]

Now: Emission & Absorption of EM radiation by atoms

## 6.5 Coupling to radiation field

Recall E & B

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad \mu = 0, 1, 2, 3$$

$$A_\mu = (-\phi, \vec{A})$$

$$x^\mu = (ct, \vec{x})$$

$$E_i = F_{i0} = -F_{0i} = -\frac{1}{c} \frac{\partial \vec{A}}{\partial t} - \frac{\partial \phi}{\partial x^i}$$

$$B_i = \frac{1}{2} \epsilon^{ijk} F_{jk} = \epsilon^{ijk} \partial_j A_k$$

E, B unchanged under gauge xforms

$$A_\mu \rightarrow A_\mu + \partial_\mu \Delta$$

For charged particle, spin  $\vec{S}$ ,

$$H = \frac{1}{2m} (\vec{p} - \frac{e}{c} \vec{A})^2 + e\phi - g_s \mu_B \frac{\vec{S}}{\hbar} \cdot (\vec{\nabla} \times \vec{A})$$

In free space (no sources) Maxwell is

$$\partial_\mu F^{\mu\nu} = 0$$

Choose Coulomb (radiation) gauge

$$A_0 = 0, \quad \vec{\nabla} \cdot \vec{A} = 0.$$

↑ transversality condition (Lorentz gauge + get rid of  $A_0$ )

Fermi (1930) showed: [see sakurai: "Advanced QM" for details]

Charged matter + EM fields can be described by [break  $A = A_{\perp} + A_{\parallel}$ ]

$$H = \underbrace{\left[ \frac{p^2}{2m} + V \right]}_{H_0} \underbrace{- \frac{e}{mc} \vec{p} \cdot \vec{A}_{\perp}}_{V(t)} + H_{\text{RAD}}^{(A_{\perp})} + \frac{e^2}{2mc^2} A_{\perp}^2 - \frac{g\mu}{\hbar} \frac{S \cdot B}{\hbar}$$

instantaneous  
↓  
Coulomb  
interaction

ignore multi-photon processes  
spin effects for now

where  $A_{\perp}$  is purely transverse field. ( $\vec{\nabla} \cdot \vec{A}_{\perp} = 0$   
6.6 Absorption cross-section ( $\vec{\nabla} \times \vec{A}_{\parallel} = 0$ )

Maxwell eqs for transverse field (drop "1"),

$$\square A^i = \left( -\frac{1}{c^2} \frac{\partial^2}{\partial t^2} + \nabla^2 \right) A^i = 0$$

Plane wave solutions

$$\vec{A} = 2A_0 \hat{\epsilon} \cos \left( \frac{\omega}{c} \hat{n} \cdot \vec{x} - \omega t \right)$$

$$\text{where } \hat{\epsilon} \cdot \hat{n} = 0$$

Energy density

$$u = \frac{1}{2} \left( \frac{E_{\text{max}}^2}{8\pi^2} + \frac{B_{\text{max}}^2}{8\pi^2} \right)$$

$$= \frac{1}{2\pi} \frac{\omega^2}{c^2} |A_0|^2$$

$$\vec{A} = A_0 \hat{\epsilon} \left[ \underbrace{e^{i(\frac{\omega}{c})\hat{n} \cdot \vec{x} - i\omega t}}_{\text{absorption}} + \underbrace{e^{-i(\frac{\omega}{c})\hat{n} \cdot \vec{x} + i\omega t}}_{\text{emission}} \right]$$