

6.3 Time-dependent perturbation theory

No analytic solution for generic $H = H_0 + V(t)$.

Must use perturbative analysis

Expand

$$C_n(t) = C_n^{(0)} + \underset{\substack{\uparrow \\ \mathcal{O}(V)}}{C_n^{(1)}(t)} + \underset{\substack{\uparrow \\ \mathcal{O}(V^2)}}{C_n^{(2)}(t)} + \dots$$

$C_n^{(0)}$ is initial state (time-independent)

Use time-evolution operator $U_I(t; t_0)$

$$|\alpha, t_0; t\rangle_I = U_I(t, t_0) |\alpha, t_0; t_0\rangle_I$$

U_I satisfies

$$i\hbar \frac{\partial}{\partial t} U_I(t, t_0) = V_I(t) U_I(t, t_0)$$

$$\text{with } U_I(t_0, t_0) = \mathbb{1}$$

$$\Rightarrow U_I(t, t_0) = \mathbb{1} - \frac{i}{\hbar} \int_{t_0}^t V_I(t') U_I(t', t_0) dt'$$

iterating

$$= \mathbb{1} - \frac{i}{\hbar} \int_{t_0}^t dt' V_I(t') \\ - \frac{i^2}{\hbar^2} \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' V_I(t') V_I(t'') + \dots$$

$$= \mathbb{1} + \sum_{n=1}^{\infty} \left(-\frac{i}{\hbar}\right)^n \int_{t_0}^t dt_1 \cdots \int_{t_0}^{t_{n-1}} dt_n V_I(t_1) V_I(t_2) \cdots V_I(t_n)$$

$$= \sum_{n=0}^{\infty} \frac{1}{n!} \left(-\frac{i}{\hbar}\right)^n \mathcal{T} \left[\int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 \cdots \int_{t_0}^{t_{n-1}} dt_n V_I(t_1) V_I(t_2) \cdots V_I(t_n) \right]$$

[Dyson series] \uparrow time-ordering operator & higher times t on left

$$U(t, t_0) = \mathcal{T} \left[e^{-\frac{i}{\hbar} \int_{t_0}^t dt' V_I(t')} \right]$$

In compact form.

Evolution of state:

Starting in state $|i\rangle$ at $t=t_0$,

$$|i, t_0; t\rangle = U_I(t, t_0) |i\rangle$$

$$= \sum_n |n\rangle \underbrace{\langle n | U_I(t, t_0) | i \rangle}_{C_n(t)}$$

since $U_I = e^{iH_0 t/\hbar} U_S e^{-iH_0 t/\hbar}$,

we have

$$|C_n(t)|^2 = |\langle n | U_I(t, t_0) | i \rangle|^2 = |\langle n | U_S(t, t_0) | i \rangle|^2$$

if $|n\rangle, |i\rangle$ are eigenvectors of H_0 .

Using $V_I(t) = e^{iH_0 t/\hbar} V(t) e^{-iH_0 t/\hbar}$

$W_{nm} = \frac{E_n - E_m}{\hbar}$

Since $U_I(t, t_0) = e^{iH_0 t/\hbar} U(t, t_0) e^{-iH_0 t/\hbar}$

between $c^{(1)}$ & $c^{(2)}$

$$|c_n|^2 = |\langle n | U_I(t, t_0) | i \rangle|^2 = |\langle n | U(t, t_0) | i \rangle|^2$$

Last time: $H = H_0 + V(t)$

$$|\psi(t)\rangle_I = \sum c_n(t) |n\rangle$$

$$c_n(t) = \langle n | U_I(t, t_0) | i \rangle$$

$|i\rangle =$ initial state

$$U_I(t, t_0) = T \left[e^{-\frac{i}{\hbar} \int_{t_0}^t dt' V_I(t')} \right]$$

pert. expansion

$$c_n(t) = c_n^{(0)} + c_n^{(1)}(t) + c_n^{(2)}(t) + \dots$$

\downarrow \downarrow
 $O(V)$ $O(V^2)$

We can expand, if initial state is $|i\rangle$,

$$\begin{aligned}
 C_n(t) &= \langle n | U_I(t, t_0) | i \rangle \\
 &= \delta_{ni} - \frac{i}{\hbar} \int_{t_0}^t dt' \langle n | V_I(t') | i \rangle \\
 &\quad - \frac{1}{\hbar^2} \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' \sum_m \langle n | V_I(t') | m \rangle \langle m | V_I(t'') | i \rangle \\
 &\quad + \dots
 \end{aligned}$$

so perturbative expansion is

$$C_n^{(0)} = \delta_{ni}$$

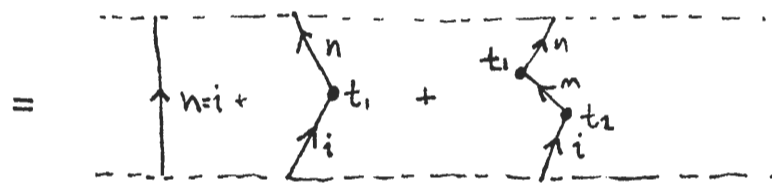
$$C_n^{(1)}(t) = -\frac{i}{\hbar} \int_{t_0}^t dt' \langle n | V_I(t') | i \rangle = -\frac{i}{\hbar} \int_{t_0}^t dt' e^{i\omega_{ni}t'} V_{ni}(t')$$

$$C_n^{(2)}(t) = -\frac{1}{\hbar^2} \sum_m \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' e^{i\omega_{nm}t' + i\omega_{mi}t''} V_{nm}(t') V_{mi}(t'')$$

↓
[trans prob - next page] ∴

Graphical depiction: "Feynman diagrams"

$$\langle n | U(t, t_0) | i \rangle = e^{-iE_n t/\hbar + iE_i t_0/\hbar} \langle n | U_I(t, t_0) | i \rangle$$



where $\begin{array}{c} t'' \\ | \\ n \\ | \\ t' \end{array} \Rightarrow e^{-iE_n(t''-t')/\hbar}$, $\begin{array}{c} t' \\ | \\ m \\ | \\ n \end{array} \Rightarrow \langle m | V(t') | n \rangle$

Transition probability $|i\rangle \rightarrow |n\rangle$, $n \neq i$, given by

$$P(i \rightarrow n) = |c_n(t)|^2 = |c_n^{(1)}(t) + c_n^{(2)}(t) + \dots|^2.$$

6.4 First order perturbation theory

1st order TDPT:

$$c_n^{(1)}(t) = -\frac{i}{\hbar} \int_{t_0}^t dt' e^{i\omega_n t'} V_{ni}(t'), \quad n \neq i$$

$$P^{(1)}(i \rightarrow n) = |c_n^{(1)}(t)|^2, \quad n \neq i$$

1st order TDPT assumes $c_n(t) = \delta_{ni}$ on RHS of Eqn for c_n 's.

Valid as long as $|c_n(t)|^2 \ll 1$, $n \neq i$

$$1 - |c_i(t)|^2 \ll 1.$$

Special cases: harmonic / constant perturbation

Assume $V(t) = \hat{V} \sin \omega t$, $t > 0$

$$V_{ni}(t) = \frac{1}{2i} \hat{V}_{ni} (e^{i\omega t} - e^{-i\omega t})$$

$$c_n^{(1)}(t) = -\frac{\hat{V}_{ni}}{2\hbar} \int_0^t e^{i\omega_n t'} [e^{i\omega t'} - e^{-i\omega t'}]$$

$$= \frac{\hat{V}_{ni}}{2\hbar i} \left[\frac{1 - e^{i(\omega_n + \omega)t}}{\omega_n + \omega} - \frac{1 - e^{i(\omega_n - \omega)t}}{\omega_n - \omega} \right]$$

If $V(t) = \hat{V} \cos \omega t$

$$C_n^{(1)} = \frac{\hat{V}_{ni}}{2\hbar} \left[\frac{1 - e^{i(\omega_{ni} + \omega)t}}{\omega_{ni} + \omega} + \frac{1 - e^{i(\omega_{ni} - \omega)t}}{\omega_{ni} - \omega} \right]$$

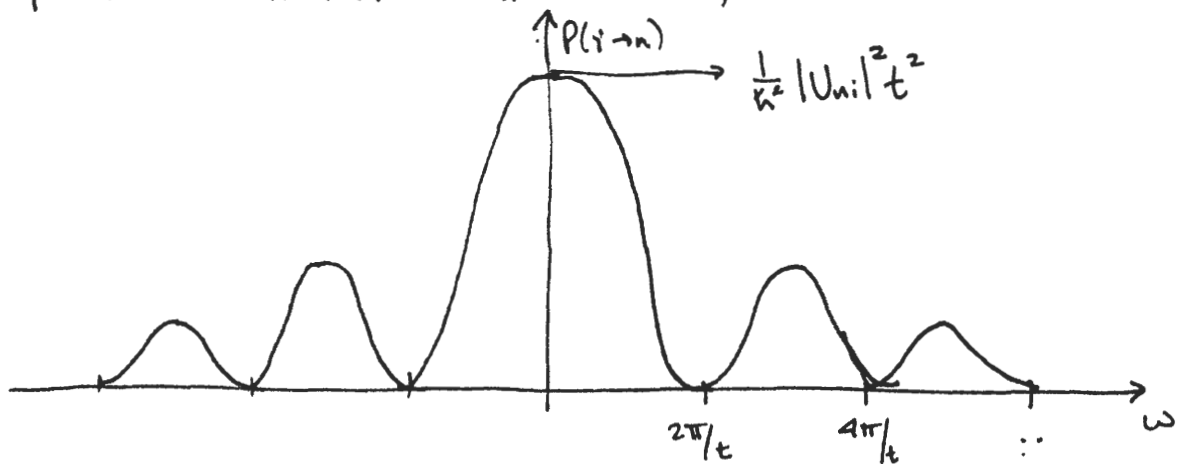
If $\omega = 0$, $V(t) = \hat{V}(t) = \text{const.}$,

$$C_n^{(1)}(t) = \frac{\hat{V}_{ni}}{\hbar \omega_{ni}} [1 - e^{i\omega_{ni}t}]$$

For constant perturbation,

$$\begin{aligned} P^{(1)}(i \rightarrow n) &= |C_n^{(1)}(t)|^2 = \frac{|\hat{V}_{ni}|^2}{(E_n - E_i)^2} [2 - 2 \cos \omega_{ni}t] \\ &= \frac{4 |\hat{V}_{ni}|^2}{(E_n - E_i)^2} \sin^2 \left[\frac{(E_n - E_i)t}{2\hbar} \right] \end{aligned}$$

Graph as function of $\omega_{ni} = (E_n - E_i)/\hbar$



[note scaling vs. $\hbar \omega_{ni}$]

For $E_n = E_i$, prob. grows as t^2 .

But — recall approx only good when $P \ll 1$.

After time t , $\Delta E \sim \frac{2\pi\hbar}{t}$.

Recalls $\Delta E \Delta t \sim \hbar$, time-energy uncertainty relation.

[Note: in completely a treatment it concerned]

Fermi's golden rule

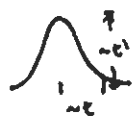
Want total transition probability

$$P^{(1)}(i \rightarrow \text{anything}) = \sum_n |C_n^{(1)}|^2.$$

When spectrum continuous (or closely spaced)
write density of states $\rho(E) dE$

$$\rho(E) = \lim_{\Delta E \rightarrow 0} \frac{(\# \text{ of states between } E - \Delta E/2, E + \Delta E/2)}{\Delta E}$$

$$\begin{aligned} P^{(1)}(i \rightarrow \text{anything}) &= \int dE_n \rho(E_n) |C_n^{(1)}|^2 \\ &= A \int \sin^2 \left[\frac{(E_n - E_i)t}{2\hbar} \right] \frac{|V_{ni}|^2}{|E_n - E_i|^2} \rho(E_n) dE_n \end{aligned}$$



For small t , Area $\sim (t^2)(t^{-1}) \sim t$.

so P goes linearly for small t , as it must.

For large t (but still small enough ω P.T. to be ok)

$$\lim_{x \rightarrow \infty} \frac{\sin^2 \alpha x}{\alpha x^2} = \pi \delta(x) \quad \left(\begin{array}{l} \int = \pi \quad \forall x \\ \lim = 0, \quad x \neq 0. \end{array} \right)$$

Transition rate: $\omega_{i \rightarrow n} = \frac{d}{dt} |c_n^{(1)}|^2$

$$\omega_{i \rightarrow n} = \frac{2\pi}{\hbar} |\hat{V}_{ni}|^2 \delta(E_n - E_i)$$

Integrating,

$$\lim_{t \rightarrow \infty} P^{(1)}(i \rightarrow \text{anything}) = \frac{2\pi}{\hbar} \overline{|\hat{V}_{ni}|^2} \rho(E_n) \Big|_{E_n \cong E_i}$$

$$\text{where } \overline{|\hat{V}_{ni}|^2} = \lim_{\Delta E \rightarrow 0} \frac{1}{\Delta E} \int_{- \Delta E/2}^{+ \Delta E/2} |\hat{V}_{ni}|^2 dE_n$$

[Valid when \hat{V}_{ni} depends smoothly on E_n
(for relevant states)]

Total transition rate = trans. prob. / unit time

$$= \frac{d}{dt} \left(\sum_n |c_n^{(1)}|^2 \right)$$

$$\omega_{i \rightarrow [n]} = \frac{2\pi}{\hbar} \overline{|\hat{V}_{ni}|^2} \rho(E_n) \Big|_{E_n \cong E_i}$$

↑
final states w/ energy $\approx E_i$

Fermi's Golden Rule

[Lecture 4]

Ura

① Problem set 2, due Monday 2/25/02

Problems 33, 35, 37, 38, 39 From Sakurai chapter 5

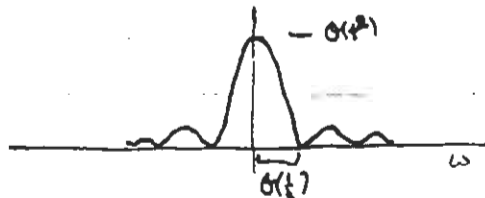
Last time:

1st order TDPT for harmonic (constant $V(t)$).

For constant $V(t) = \hat{V}$, $C_n^{(1)} = \frac{\hat{V}_{ni}}{\hbar \omega_{ni}} [1 - e^{-i\omega_{ni}t}]$

$$P^{(1)}(i \rightarrow n) = \frac{4 |\hat{V}_{ni}|^2}{(\hbar \omega_{ni})^2} \text{Si}^2 \left[\frac{(E_n - E_i)t}{2\hbar} \right]$$

$$= |C_n^{(1)}(t)|^2$$



$$\lim_{t \rightarrow \infty} \sim t \pi \delta(\omega)$$

~~$$\lim_{x \rightarrow \infty} \frac{\text{Si}^2(x)}{x^2} = \pi \delta(x)$$~~

$$\text{Transition rate } W_{i \rightarrow n} = \frac{d}{dt} |C_n^{(1)}|^2 = \frac{2\pi}{\hbar} |\hat{V}_{ni}|^2 \delta(E_n - E_i)$$

(t big enough for EC, short enough, TDPT ok.)

Introduce density of states $\rho(E)$

$$W_{i \rightarrow n} = \frac{2\pi}{\hbar} |\hat{V}_{ni}|^2 \rho(E_n) \Big|_{E_n = E_i}$$

↑
final states w/ $E_n = E_i$

Fermi's Golden Rule

$$\left[|\hat{V}|^2 = \lim_{\Delta E \rightarrow 0} \frac{1}{\Delta E} \int_{-0\epsilon/2}^{0\epsilon/2} |\hat{V}_{ni}|^2 dE_n \quad \text{if } \rho, \hat{V} \text{ smooth on } \text{final states} \right]$$

Back to harmonic perturbation

$$V(t) = V e^{i\omega t} + V^+ e^{-i\omega t}$$

$$C_n^{(1)} = \frac{1}{\hbar} \left[\underbrace{\frac{1 - e^{i(\omega_{ni} + \omega)t}}{\omega_{ni} + \omega}}_{\text{peaked near } \omega = -\omega_{ni}} V_{ni} + \underbrace{\frac{1 - e^{i(\omega_{ni} - \omega)t}}{\omega_{ni} - \omega}}_{\text{peaked near } \omega = \omega_{ni}} V_{ni}^+ \right]$$

$\omega \equiv -\omega_{ni}$: Stimulated emission

$$|C_n^{(1)}|^2 \approx \frac{4 |V_{ni}|^2}{\hbar^2 (\omega + \omega_{ni})^2} \sin^2 \left[(\omega + \omega_{ni})^2 t / 2 \right]$$

Transition rate \rightarrow state w/ energy E_n at large t

$$\omega_{i \rightarrow n} = \frac{2\pi}{\hbar} |V_{ni}|^2 \delta(E_n - E_i + \hbar\omega)$$

$$\omega_{i \rightarrow [n]} = \frac{2\pi}{\hbar} \overline{|V_{ni}|^2} \rho(E_n) \Big|_{E_n \approx E_i - \hbar\omega}$$

total emission rate 

$\omega \equiv \omega_{ni}$: absorption

$$|C_n^{(1)}|^2 \approx \frac{4 |V_{ni}^+|^2}{\hbar^2 (\omega - \omega_{ni})^2} \sin^2 \left[(\omega - \omega_{ni})^2 t / 2 \right]$$

$$\omega_{i \rightarrow n} = \frac{2\pi}{\hbar} |V_{ni}^+|^2 \delta(E_n - E_i - \hbar\omega)$$

$$\omega_{i \rightarrow [E]} = \frac{2\pi}{\hbar} |V_{ni}^+|^2 \rho(E_n) \Big|_{E_n = E_i + \hbar\omega}$$

total absorption rate



So - harmonic perturbation causes stimulated emission or absorption in units of $\hbar\omega$.

- Just what we expect if background made up of quanta w/ $E = \hbar\omega$!

For transitions to occur & satisfy energy conservation, must have

(a) final states exist over continuous energy range, to match $\Delta E = \hbar\omega$ for fixed perturbation frequency ω
-or-

(b) Perturbation must cover sufficiently wide spectrum of ω so that discrete transition with a fixed $\Delta E = \hbar\omega$ is possible.

- Note that spectral lines are not really sharp, due to decay processes.

Note: For two discrete states, $\omega_{i \rightarrow n}^{(ab)} = \omega_{n \rightarrow i}^{(em)}$ in semiclassical calc.
since $|V_{ni}|^2 = |V_{ni}^+|^2$

\Rightarrow Detailed balance

[Really, only true @ $T = \infty$ when rad. field quantized]

Now: Emission & Absorption of EM radiation by atoms

6.5 Coupling to radiation field

Recall E & M

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu \quad \mu = 0, 1, 2, 3$$

$$A_\mu = (-\phi, \vec{A})$$

$$x^\mu = (ct, \vec{x})$$

$$E_i = F_{i0} = -F_{0i} = -\frac{1}{c} \frac{\partial \vec{A}}{\partial t} - \frac{\partial \phi}{\partial x^i}$$

$$B^i = \frac{1}{2} \epsilon^{ijk} F_{jk} = \epsilon^{ijk} \partial_j A_k$$

E, B unchanged under gauge xforms

$$A_\mu \rightarrow A_\mu + \partial_\mu \Delta$$

For charged particle, spin \vec{S} ,

$$H = \frac{1}{2m} (\vec{p} - \frac{e}{c} \vec{A})^2 + e\phi - g_s \mu_B \frac{\vec{S}}{\hbar} \cdot (\nabla \times \vec{A})$$

In free space (no sources) Maxwell is

$$\partial_\mu F^{\mu\nu} = 0$$

Choose Coulomb (radiation) gauge

$$A_0 = 0, \quad \vec{\nabla} \cdot \vec{A} = 0.$$

↑ transversality condition (Lorentz gauge + get rid of A_0)

Fermi (1970) showed: [see Sakurai: "Advanced QM" for details]

Charged matter + EM fields can be described by [break $A = A_{\perp} + A_{\parallel}$]

$$H = \underbrace{\left[\frac{p^2}{2m} + V \right]}_{H_0} \underbrace{- \frac{e}{mc} \vec{p} \cdot \vec{A}_{\perp}}_{V(t)} + H_{\text{RAD}}^{(A_{\perp})} + \frac{e^2}{2mc^2} A_{\perp}^2 = \frac{q\mu}{\hbar} \frac{S \cdot B}{\cdot B}$$

instantaneous
Coulomb
interaction

ignore multi-photon
spin effects for now

where A_{\perp} is purely transverse field. ($\vec{\nabla} \cdot \vec{A}_{\perp} = 0$)

6.6 Absorption cross-section

$$\vec{\nabla} \times \vec{A}_{\parallel} = 0$$

Maxwell eqs for transverse field (drop "1"),

$$\square A^i = \left(-\frac{1}{c^2} \frac{\partial^2}{\partial t^2} + \nabla^2 \right) A^i = 0$$

Plane wave solutions

$$\vec{A} = 2A_0 \hat{\epsilon} \cos \left(\frac{\omega}{c} \hat{n} \cdot \vec{x} - \omega t \right)$$

$$\text{where } \hat{\epsilon} \cdot \hat{n} = 0$$

Energy density

$$u = \frac{1}{2} \left(\frac{E_{\text{max}}^2}{8\pi^2} + \frac{B_{\text{max}}^2}{8\pi^2} \right)$$

$$= \frac{1}{2\pi} \frac{\omega^2}{c^2} |A_0|^2$$

$$\vec{A} = A_0 \hat{\epsilon} \left[\underbrace{e^{i(\frac{\omega}{c})\hat{n} \cdot \vec{x} - i\omega t}}_{\text{absorption}} + \underbrace{e^{-i(\frac{\omega}{c})\hat{n} \cdot \vec{x} + i\omega t}}_{\text{emission}} \right]$$